

Estimating the H₂ column density using combined molecular line intensities in the Orion B cloud

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Science Question

Why do we study Giant Molecular Clouds (GMCs) ?

To investigate the process of star formation in our galaxy and elsewhere

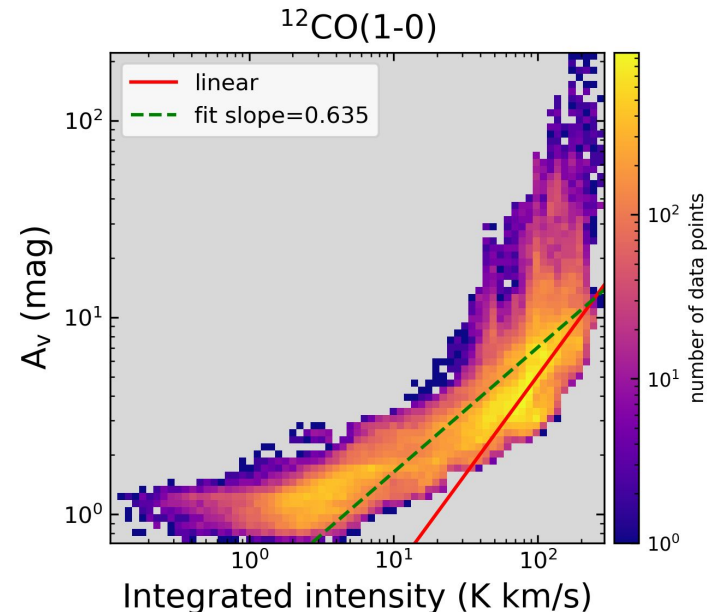
How?

- Cold H₂ is invisible → **dust** and/or **molecular lines** (CO, HCO+, ... isotopologues)
 - **dust** : optically thin, but for **SED** need **FIR** (low angular resolution), lacks velocity information
 - **rotational lines** : velocity information, achieve high angular resolution from ground (e.g., X_{CO})

Can we go beyond X_{CO} and estimate N_{H_2} with more lines and machine learning techniques?



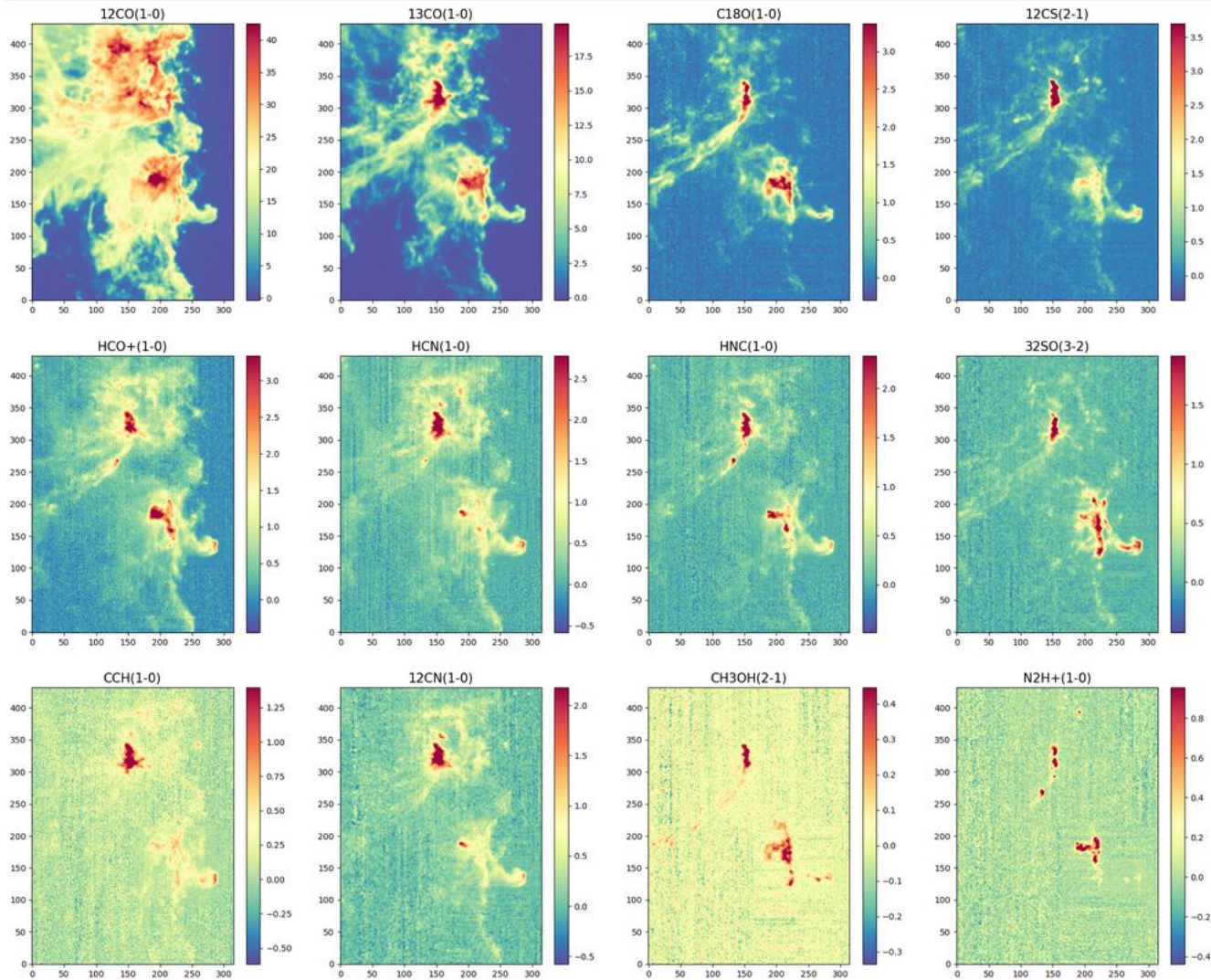
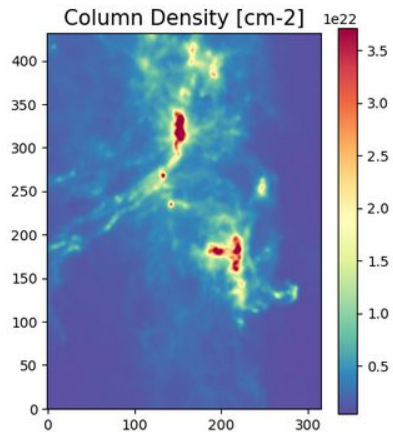
IRAM 30m Telescope , K. Zacher



The ORION-B Dataset

IRAM 30-m Telescope

- Resolution 0.07 pc/px
- Image Size is 5x7 pc
- 12 Millimeter Rotational Transition Lines



Correlation between different line intensities

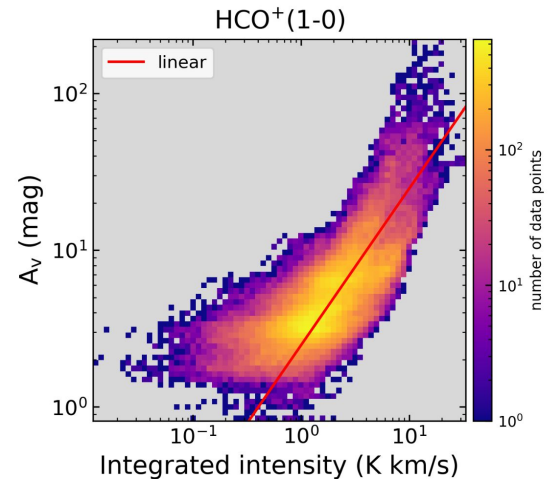
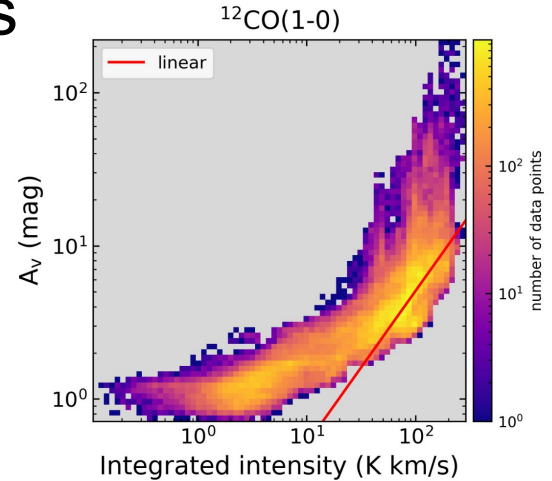
Not only CO molecular-line intensity but also other molecular-line intensities have relations to $N(\text{H}_2)$

We want to explore

- if different lines correlate to each other
- if the combination of lines has a connection to $N(\text{H}_2)$

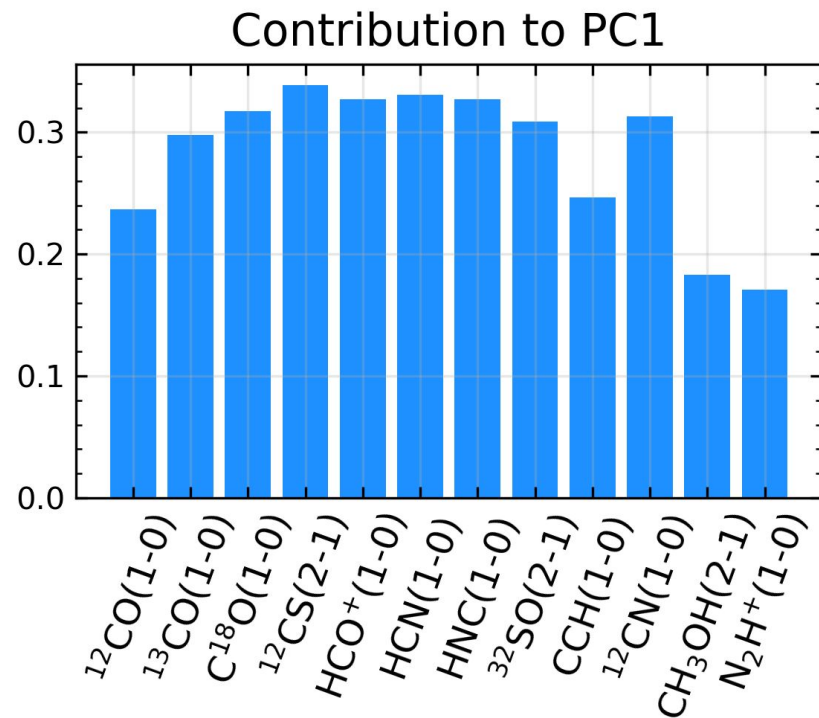
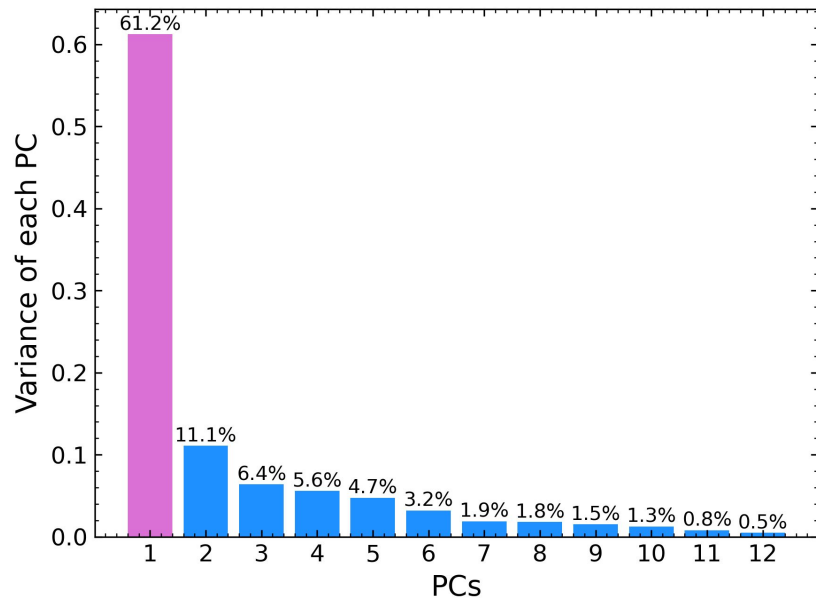
by performing Principal Component Analysis (PCA)
on the data of all 12 line intensities

- find linear correlations in the data
- useful for high dimensional data space (e.g., 12 lines)

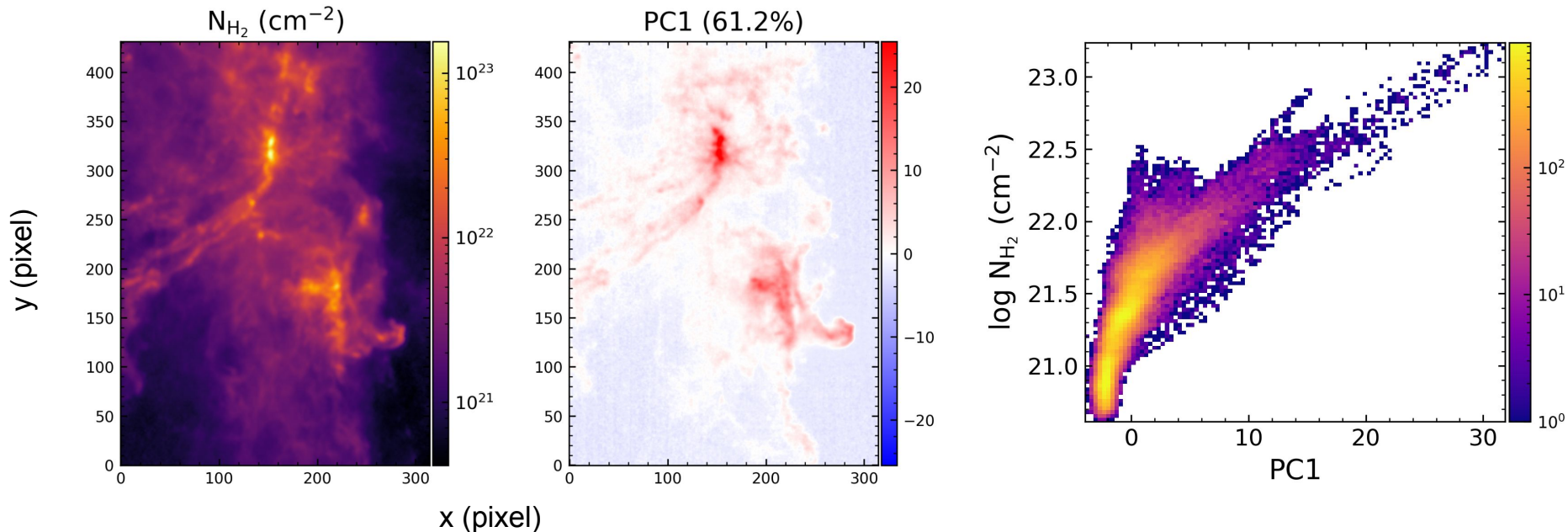


Which component is the most important?

PCA finds intrinsic characteristics from the data (PC-space) and shows which characteristics are more useful or less useful to explain the distribution of data.

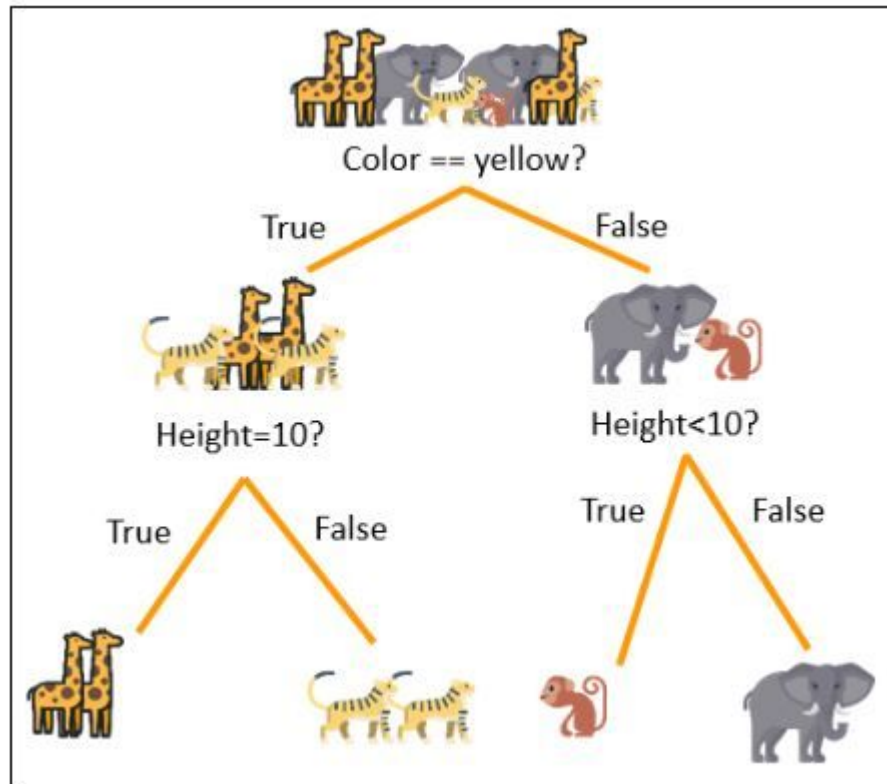
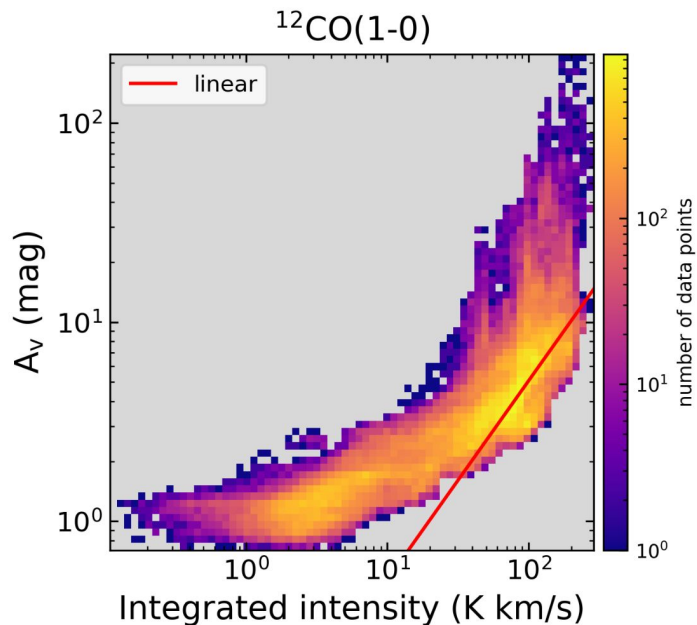


Relation between PC1 and $N(\text{H}_2)$



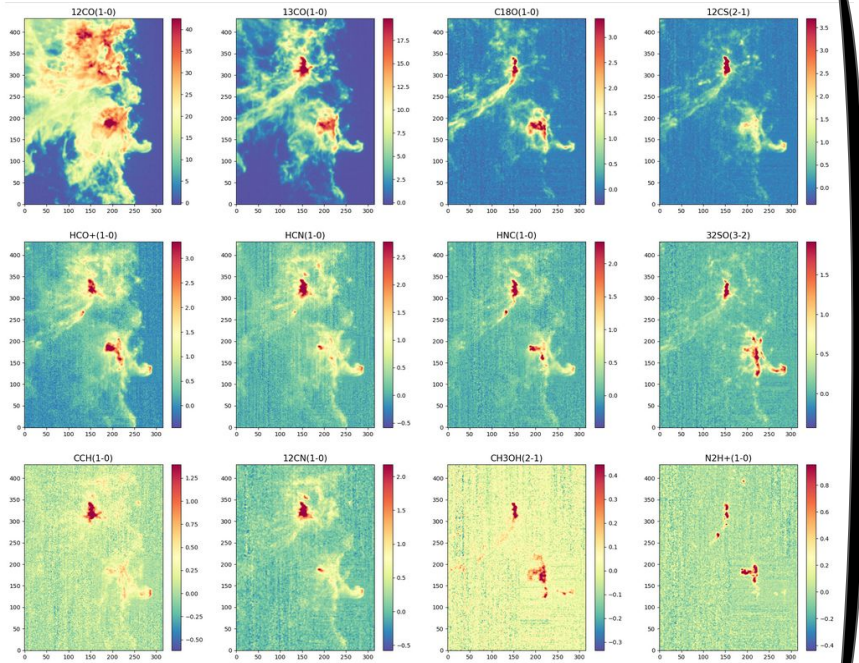
- All line intensities are related to $N(\text{H}_2)$
- Not a perfect linear correlation.
- We need a non-linear function to link the line intensities and $N(\text{H}_2)$

How do we “learn” $N(\text{H}_2)$ from line intensities?

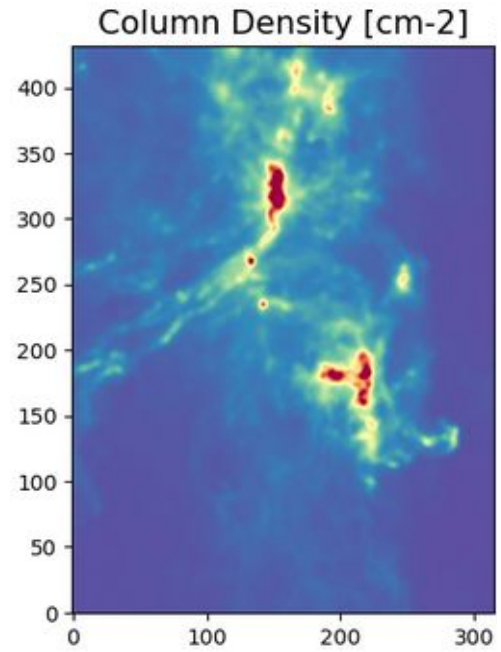


How do we “learn” $N(\text{H}_2)$ from line intensities?

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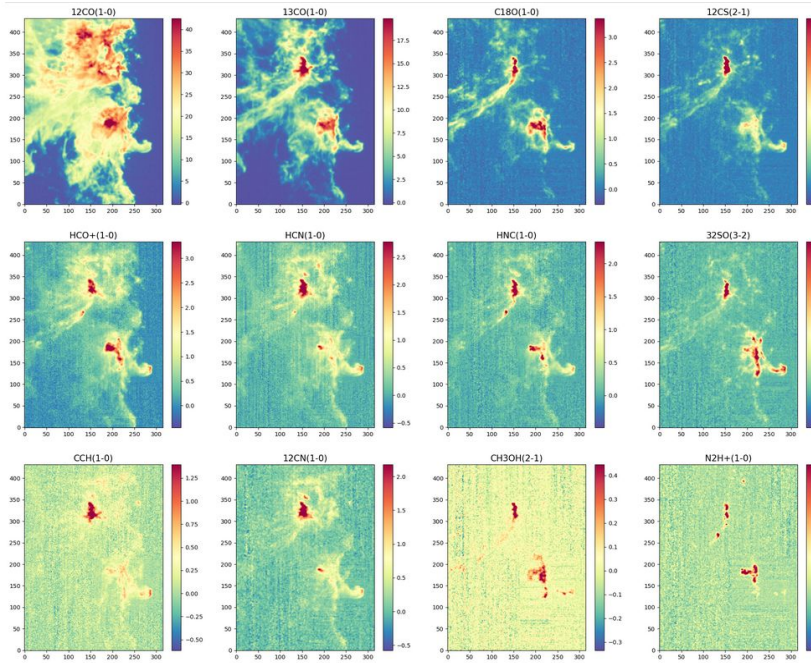


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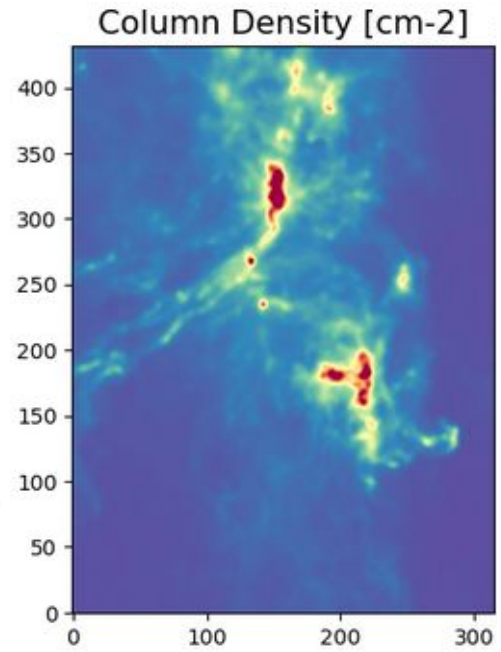
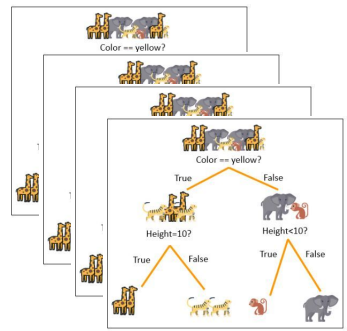


How do we “learn” $N(\text{H}_2)$ from line intensities?

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Use Random Forests!

- ★ Collection of decision trees
- ★ Good for non-linearly correlated data,
- ★ No need to normalize data, remove blank or missing values etc.
- ★ Randomizes the input sample and averages the result

How is the dataset prepared?

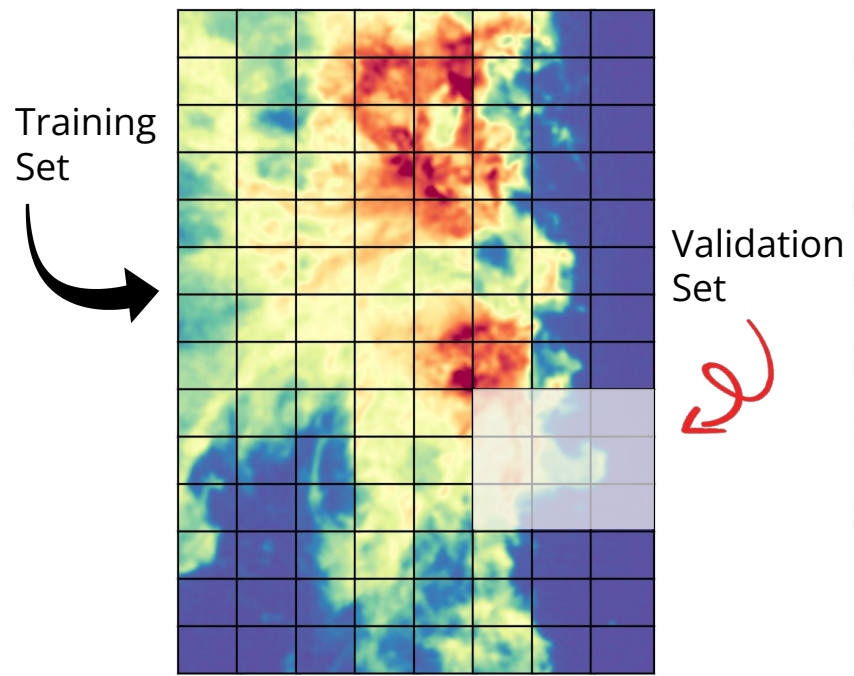
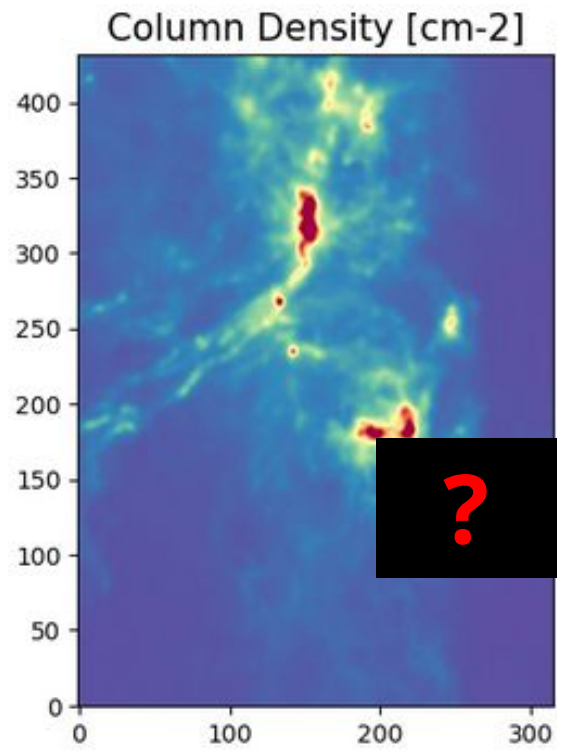


Illustration: Splitting the ^{12}CO (1-0) dataset

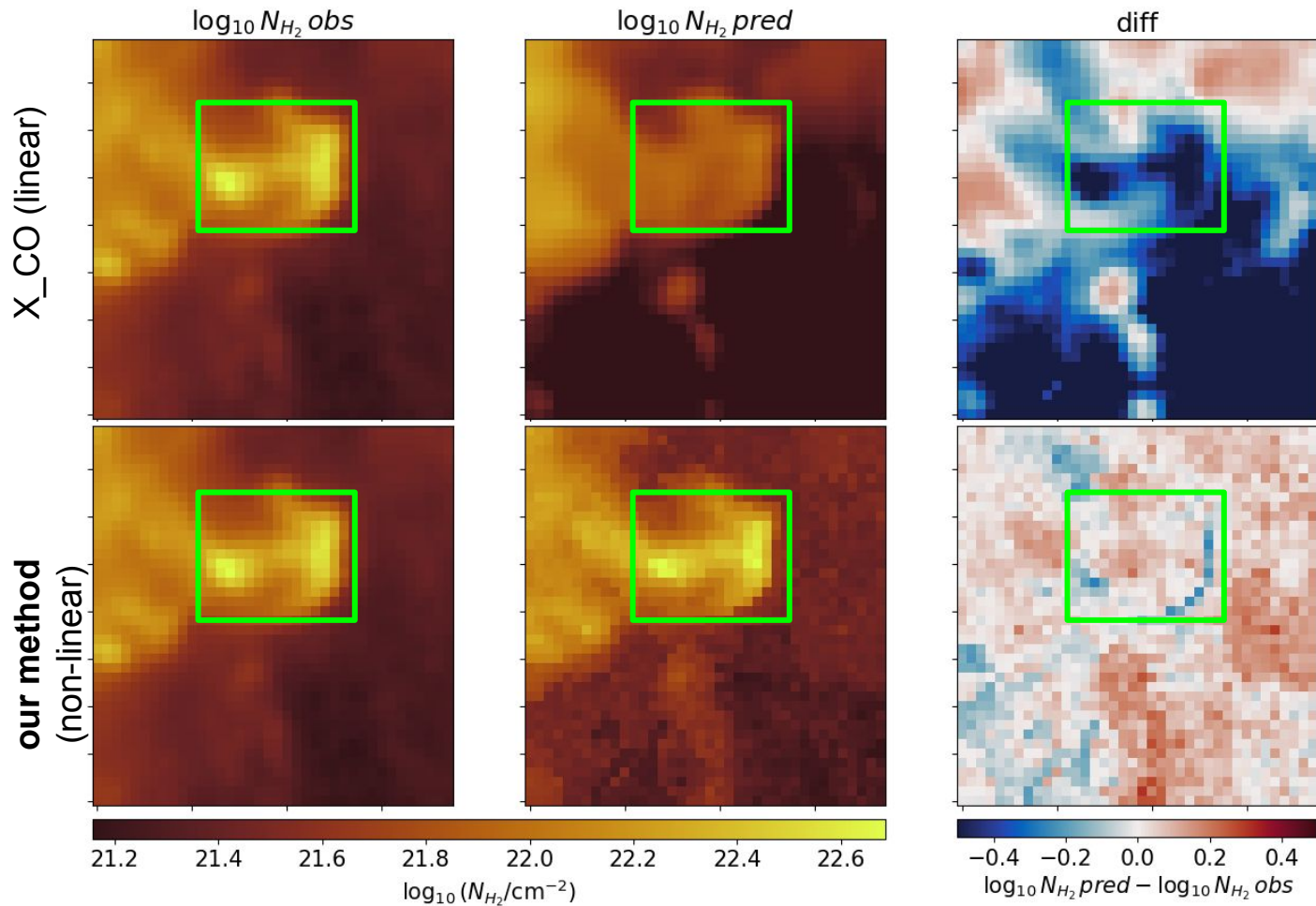


- ★ Split the correlation data into Training and Validation sets
- ★ The Random Forest regression algorithm **predicts the column densities of the validation region**

Validation



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Summary and Perspectives

What have we done?

- ★ Investigate the relation between molecular lines and the $N(\text{H}_2)$
- ★ Perform PCA to find correlations between lines
- ★ Perform RF to link molecular lines to $N(\text{H}_2)$ non-linearly

What did we find?

- ★ Our RF model works better than simply using X_{CO}

What will we do next?

- ★ Test the model robustness on **other data**
 - **well known** GMCs → check on the interpretability of the results
- ★ Apply the model to **new data**



International Summer School on the ISM of Galaxies.

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PCA methodology

- ★ *unsupervised learning*
- ★ *transforms the data to a new coordinate system*
- ★ *the **greatest variance** by some scalar projection of the data comes to lie on the first coordinate (called the first principal component), the second greatest variance on the second coordinate, and so on.*

